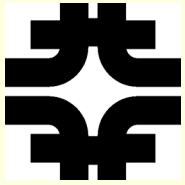


Future Neutrino Experiments

Gina Rameika

Annual Users Meeting

June 6, 2007



Outline

- The Neutrino Landscape
- Fermilab neutrino facilities
- Evolution of the current program
- Planning for the future
 - Long baseline options
 - R & D along the way



The Neutrino Landscape

Three neutrino flavors are related to three neutrino mass states by a mixing matrix :

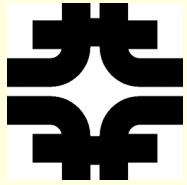
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{bmatrix} c\theta_{12} & s\theta_{12} & \\ -s\theta_{12} & c\theta_{12} & \\ & & 1 \end{bmatrix} \begin{bmatrix} c\theta_{13} & s\theta_{13}e^{-i\delta} \\ & 1 \\ -s\theta_{13}e^{i\delta} & c\theta_{13} \end{bmatrix} \begin{bmatrix} 1 & & \\ & c\theta_{23} & s\theta_{23} \\ & -s\theta_{23} & c\theta_{23} \end{bmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

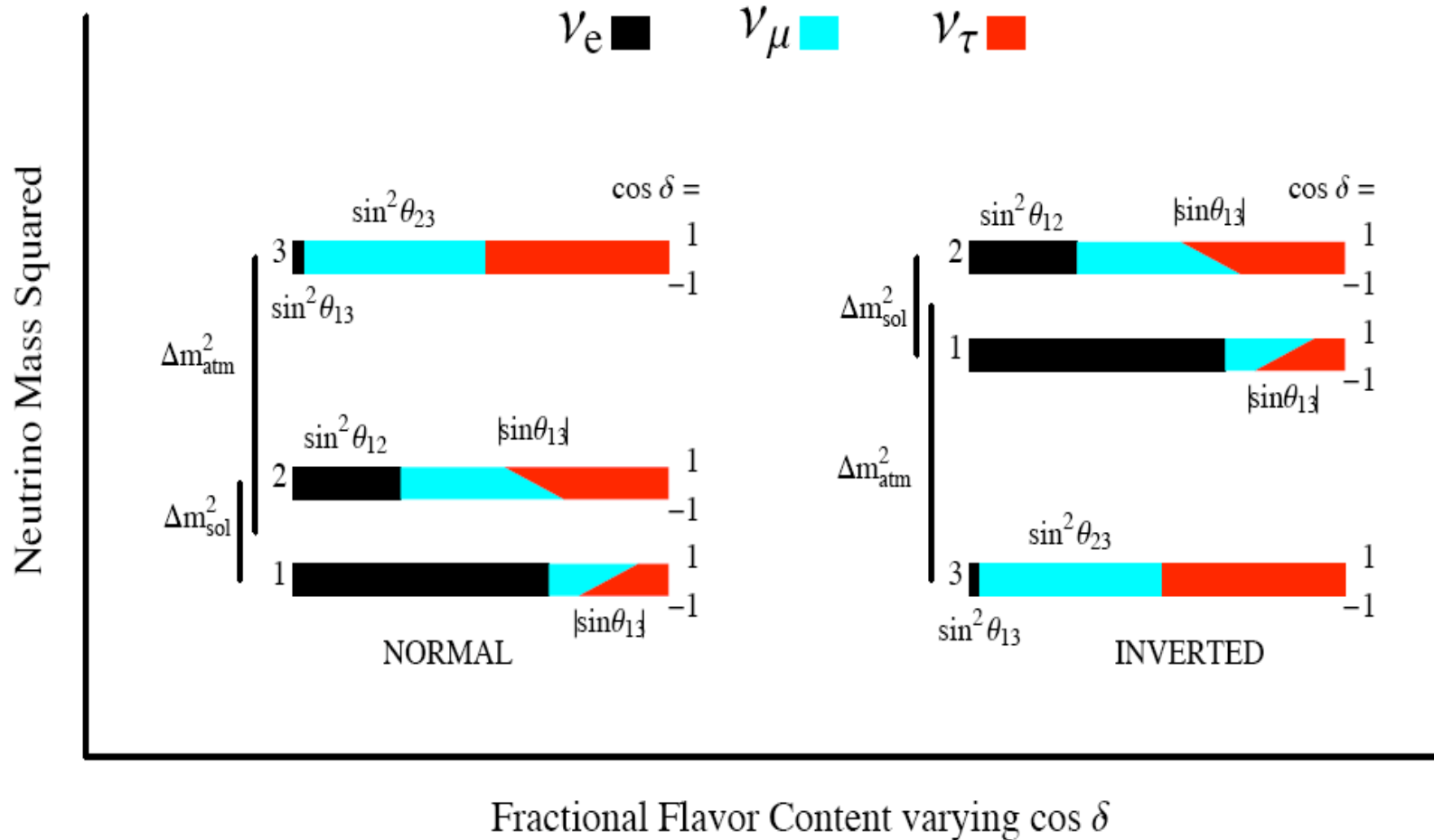
$$P(\nu_e \rightarrow \nu_\mu) = \sin^2\theta_{12} \sin^2(1.27\Delta m_{12}^2 L/E)$$

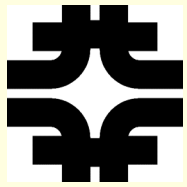
$$P(\nu_\mu \rightarrow \nu_\tau) = \sin^2\theta_{23} \sin^2(1.27\Delta m_{23}^2 L/E)$$

$$P(\nu_\mu \rightarrow \nu_e) \approx \sin^2\theta_{23} \sin^2 2\theta_{13} \sin^2(1.27\Delta m_{31}^2 L/E)$$



The Neutrino Landscape





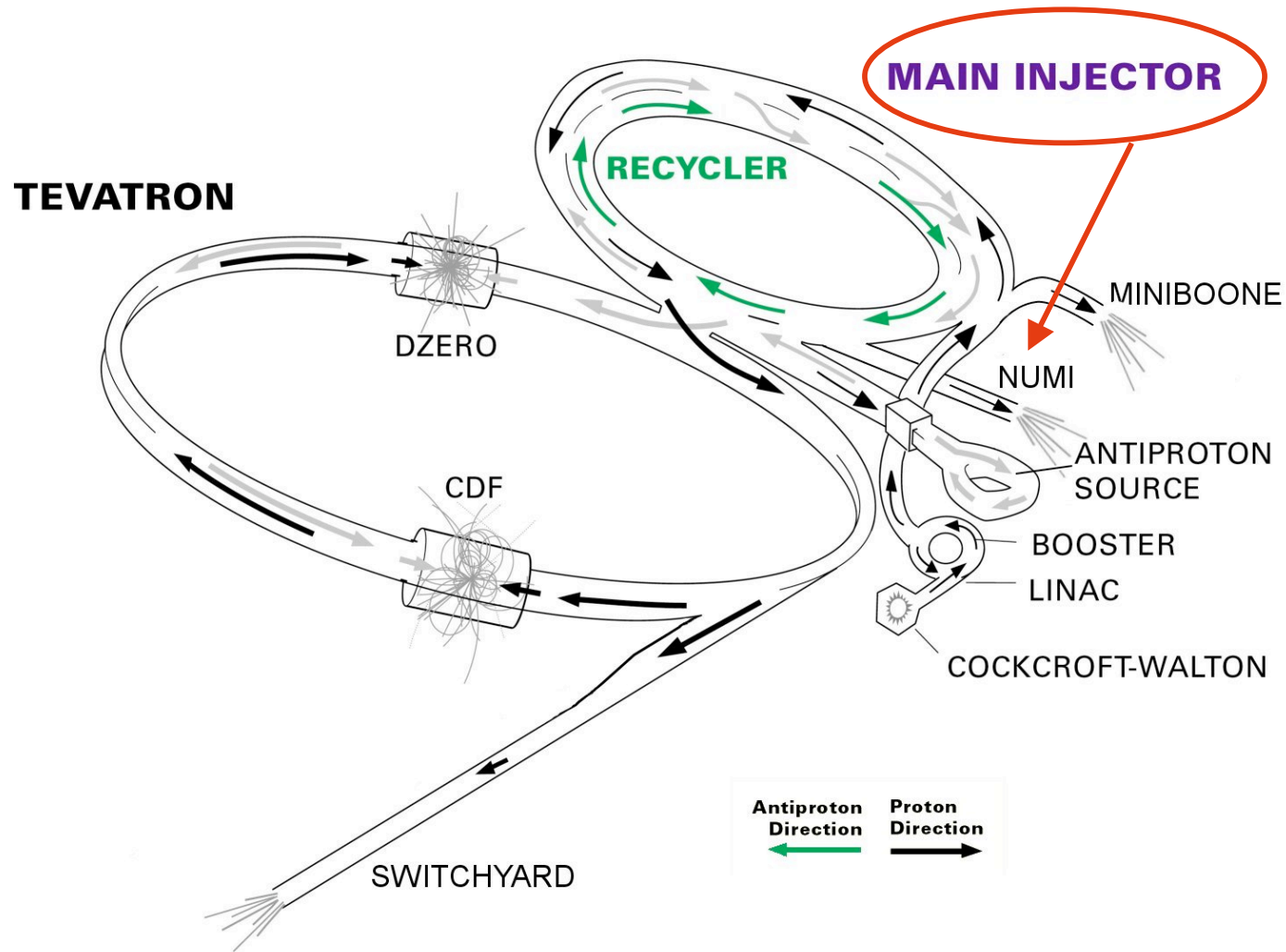
The Neutrino Landscape

- The Open Questions
 - What is the value of $\sin^2 2\theta_{13}$?
 - Just below the current limit? ~ 0.14
 - Very, very small? < 0.01
 - Exactly zero?
 - Is the mass hierarchy normal or inverted?
 - Most theoretical models for neutrino mass favor the normal hierarchy, but as experimentalists, we must measure it..
 - Is CP violated in the neutrino sector?
 - Can we determine that δ_{CP} is different from 0 or π ?
 - Are the anti-neutrino transitions different than the neutrinos?



Fermilab neutrino facilities

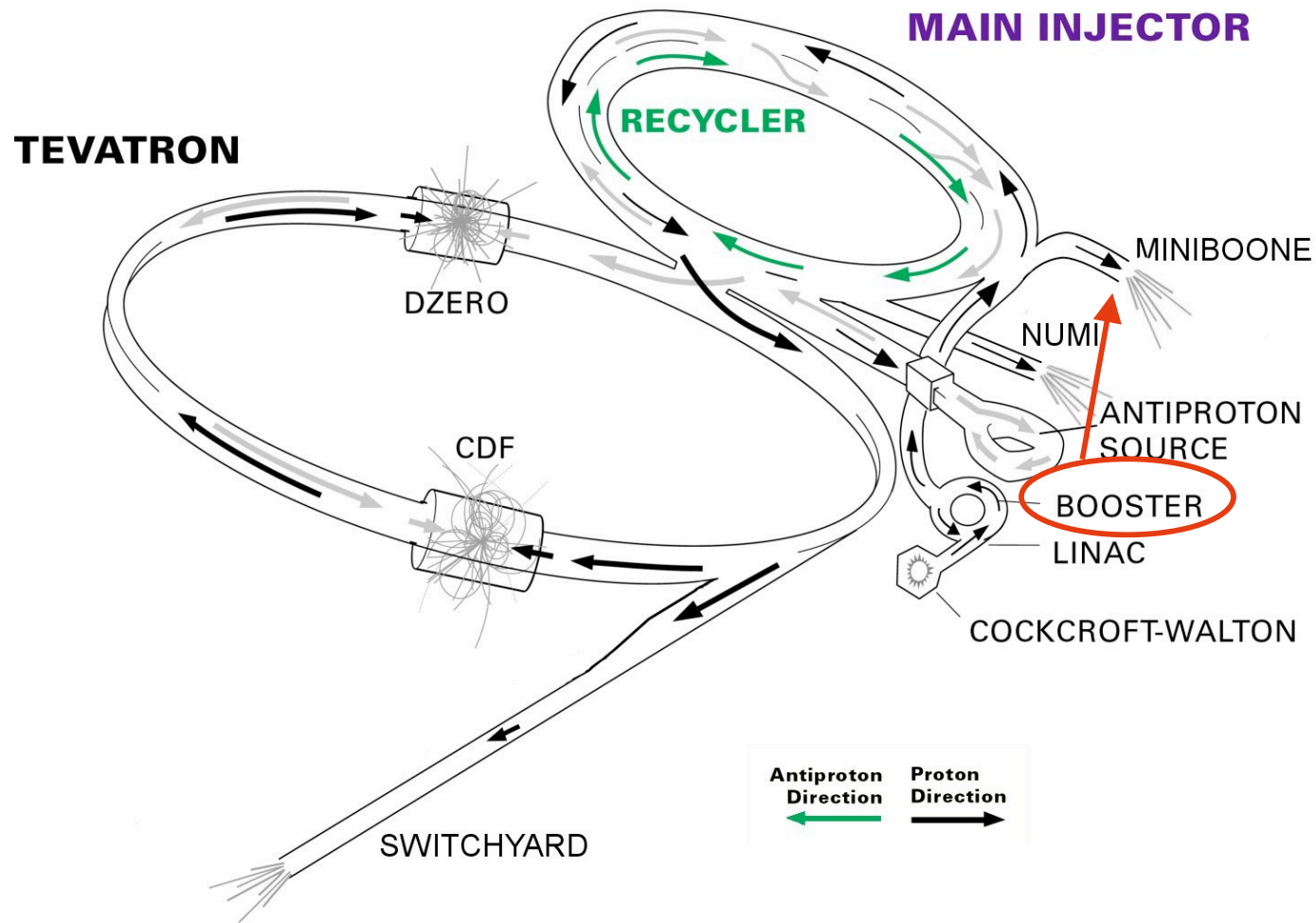
FERMILAB'S ACCELERATOR CHAIN

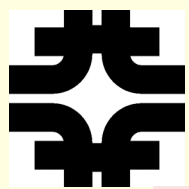




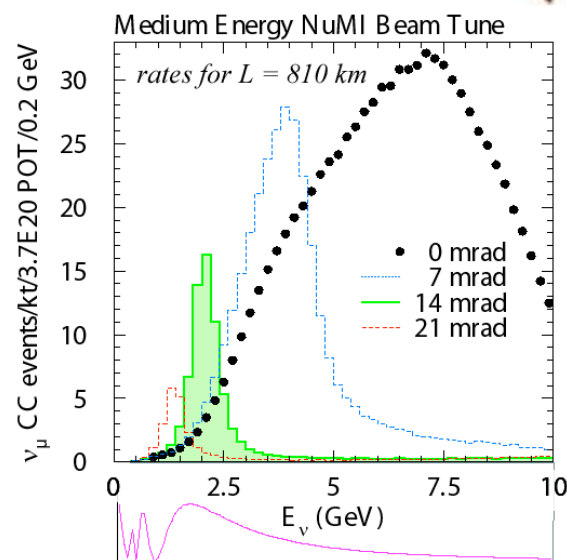
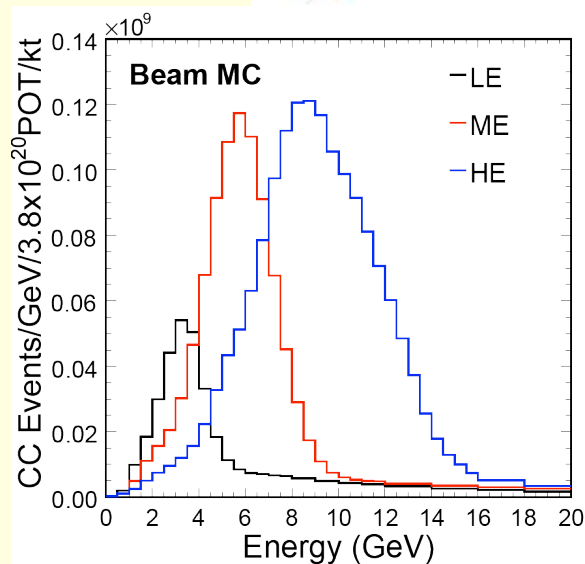
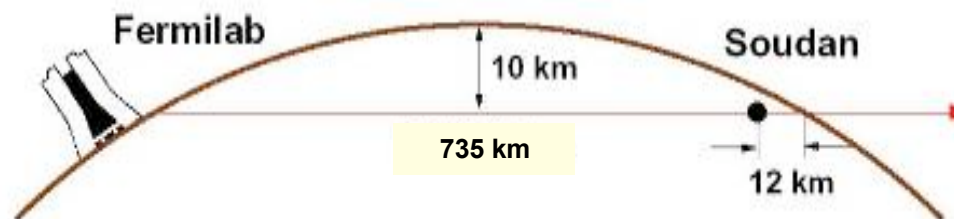
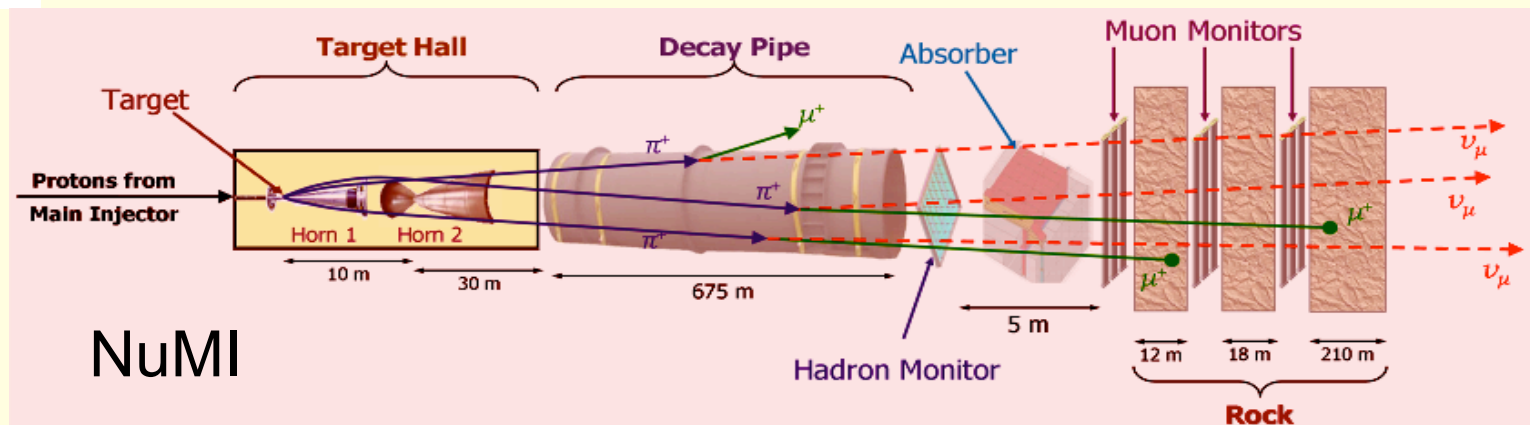
Fermilab neutrino facilities

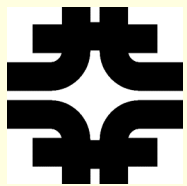
FERMILAB'S ACCELERATOR CHAIN





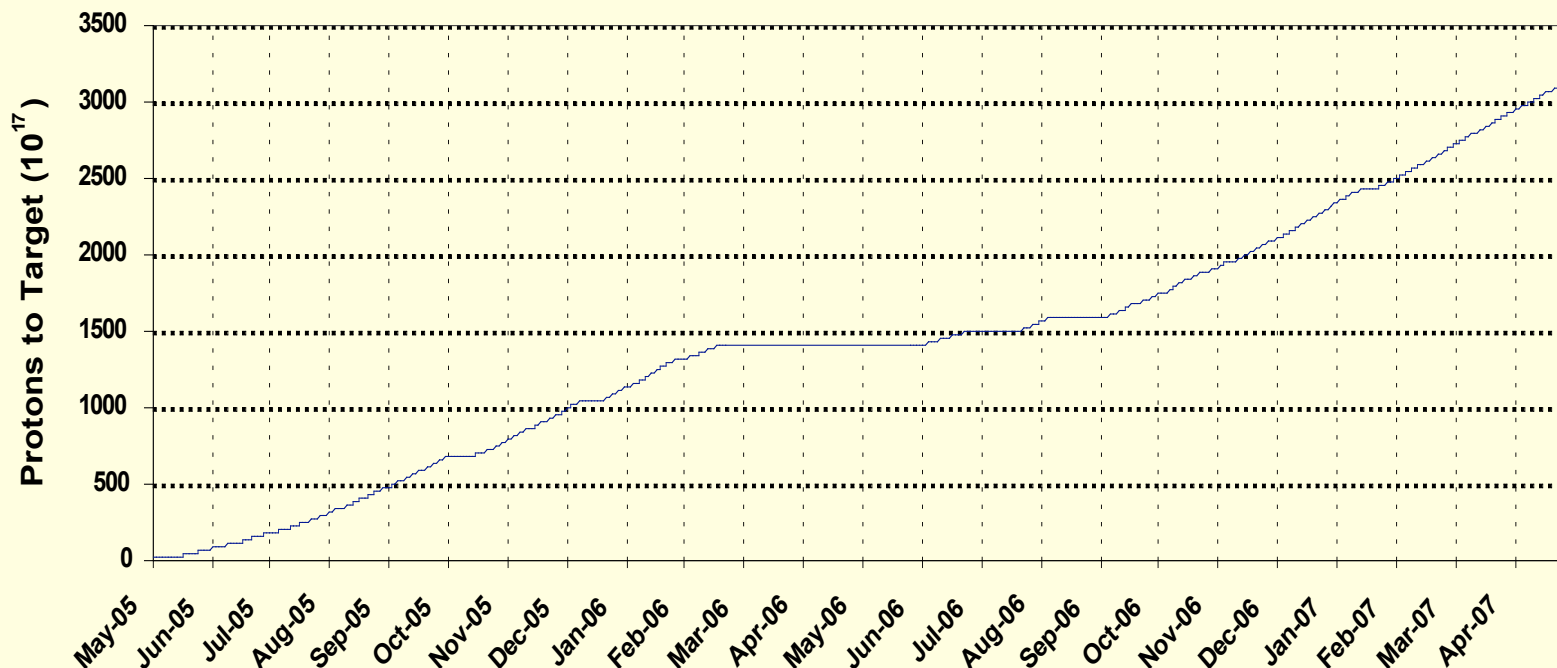
The NuMI Neutrino Beam





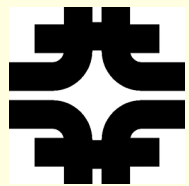
NuMI Operation

Total NuMI Protons

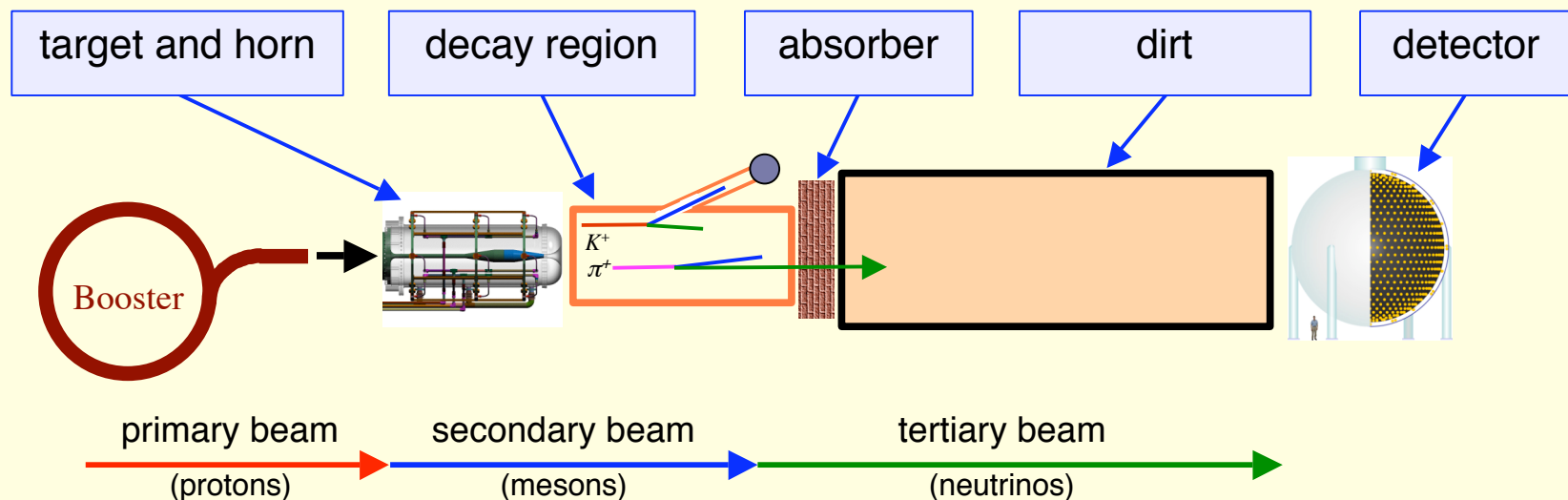


>3e20
pot

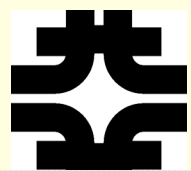
- Recent Achievements:
 - $> 4 \times 10^{13}$ protons in a pulse (exceeds design intensity)
 - 325 kW beam power



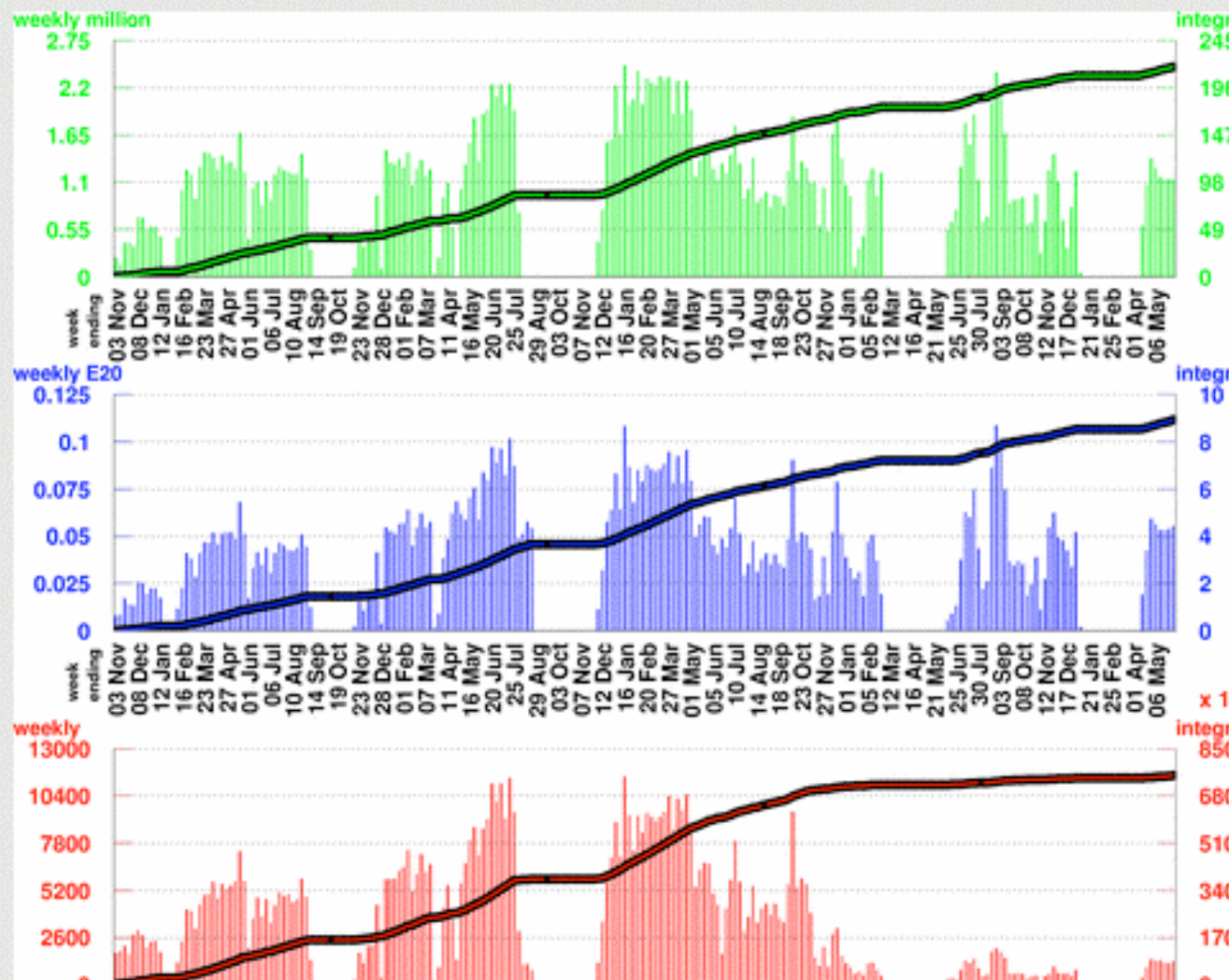
The Booster Neutrino Beam



- 4-5e12 ppp; average cycle time 2Hz
 - Last week achieved record average intensity of 4.9e12 ppp
- Operates in neutrino and anti-neutrino mode



The Booster Neutrino Beam



Number of Horn Pulses

To date: 217.47 million

Largest week: 2.46 million

Latest week: 1.13 million

Number of Protons on Target

To date: 8.9293 E20

Largest week: 0.1085 E20

Latest week: 0.0554 E20

Number of Neutrino Events

To date: 753412

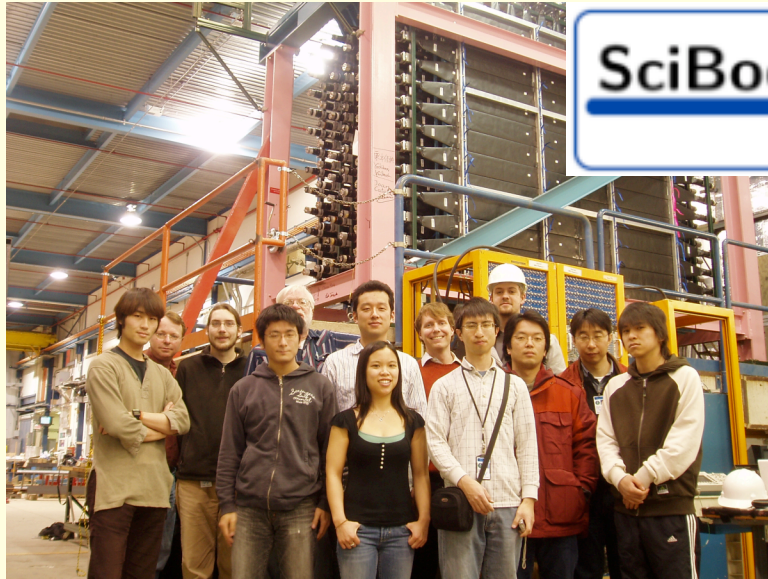
Largest week: 11447

Latest week: 1306

Operation since Nov. 2002



The Nu Experiments



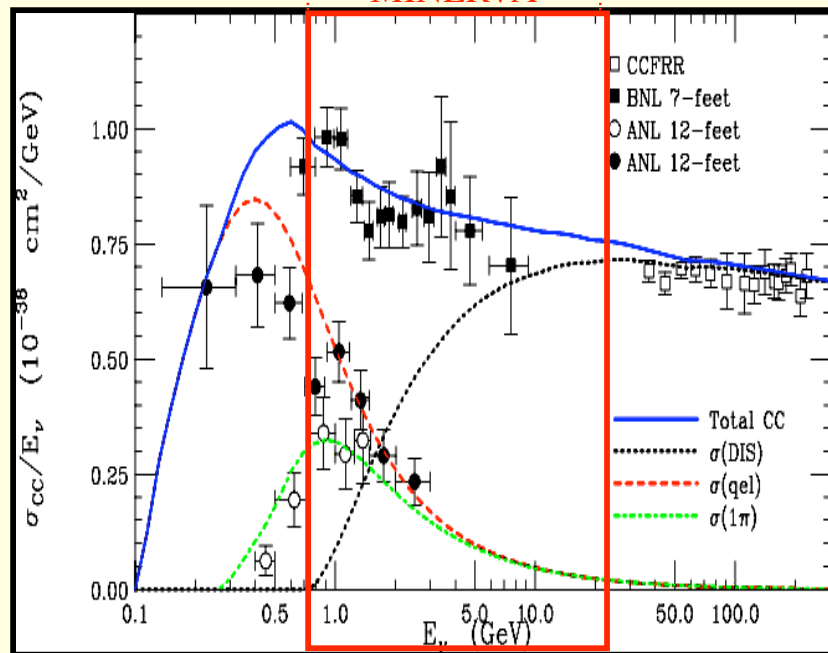
MINERvA



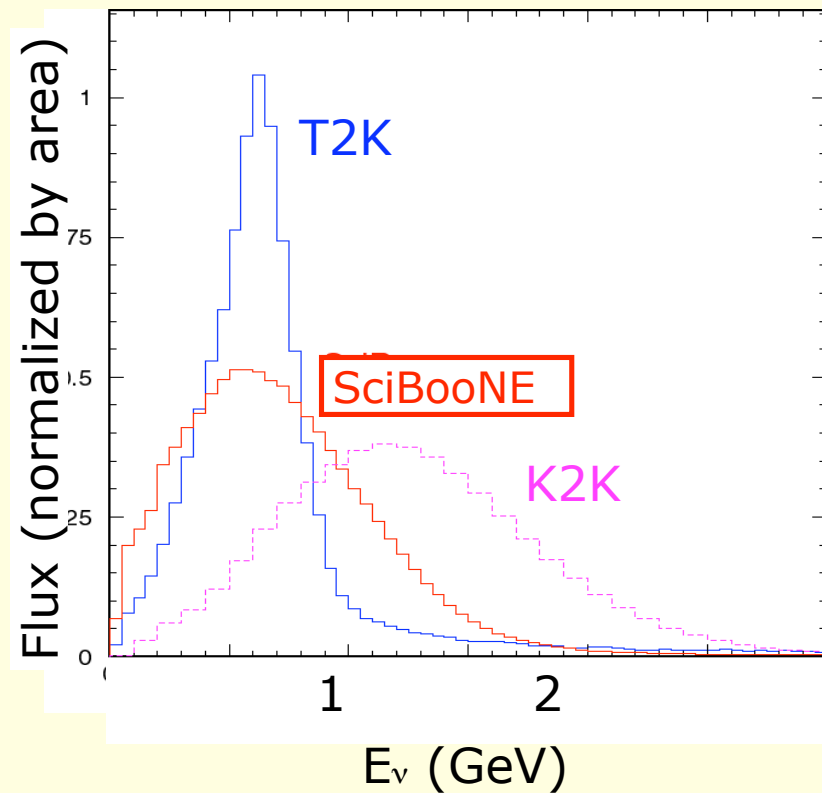


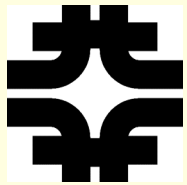
- The MINERvA and SciBooNE experiments will provide much needed data on neutrino cross sections that will help improve future measurements of the mass-mixing parameters.

Energy range of
NuMI and
MINERvA



T2K
NOvA
MINOS

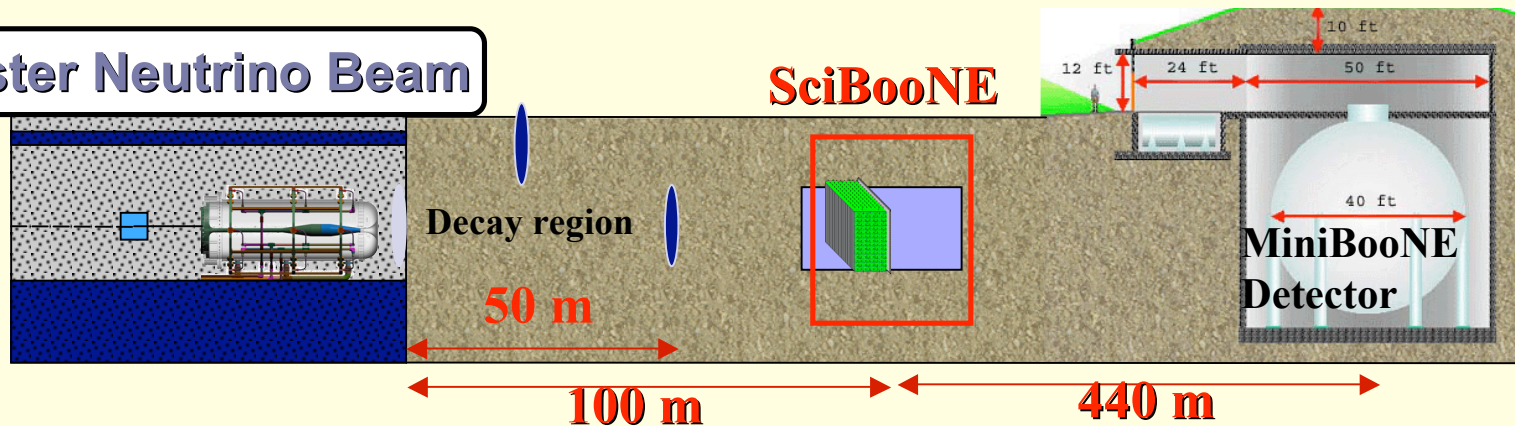




Status of SciBooNE



Booster Neutrino Beam



- **2005 Dec - Proposal to FNAL PAC**
- 2006 Mar - First collaboration meeting
- 2006 Apr - Begin work on MRD
- 2006 Jul - Detectors arrive from Japan
- 2006 Sep - Groundbreaking for detector hall
- 2006 Nov - EC assembly complete
- 2006 Jan - Civil construction complete
- 2007 Feb - SciBar assembly complete
- 2007 Mar - MRD assembly complete
- 2007 Mar - Detector hall ready
- 2007 Mar - Cosmic ray data
- 2007 Apr - Detector installation
- 2007 May - Detector commissioning
- **2007 June - Beam data run!**

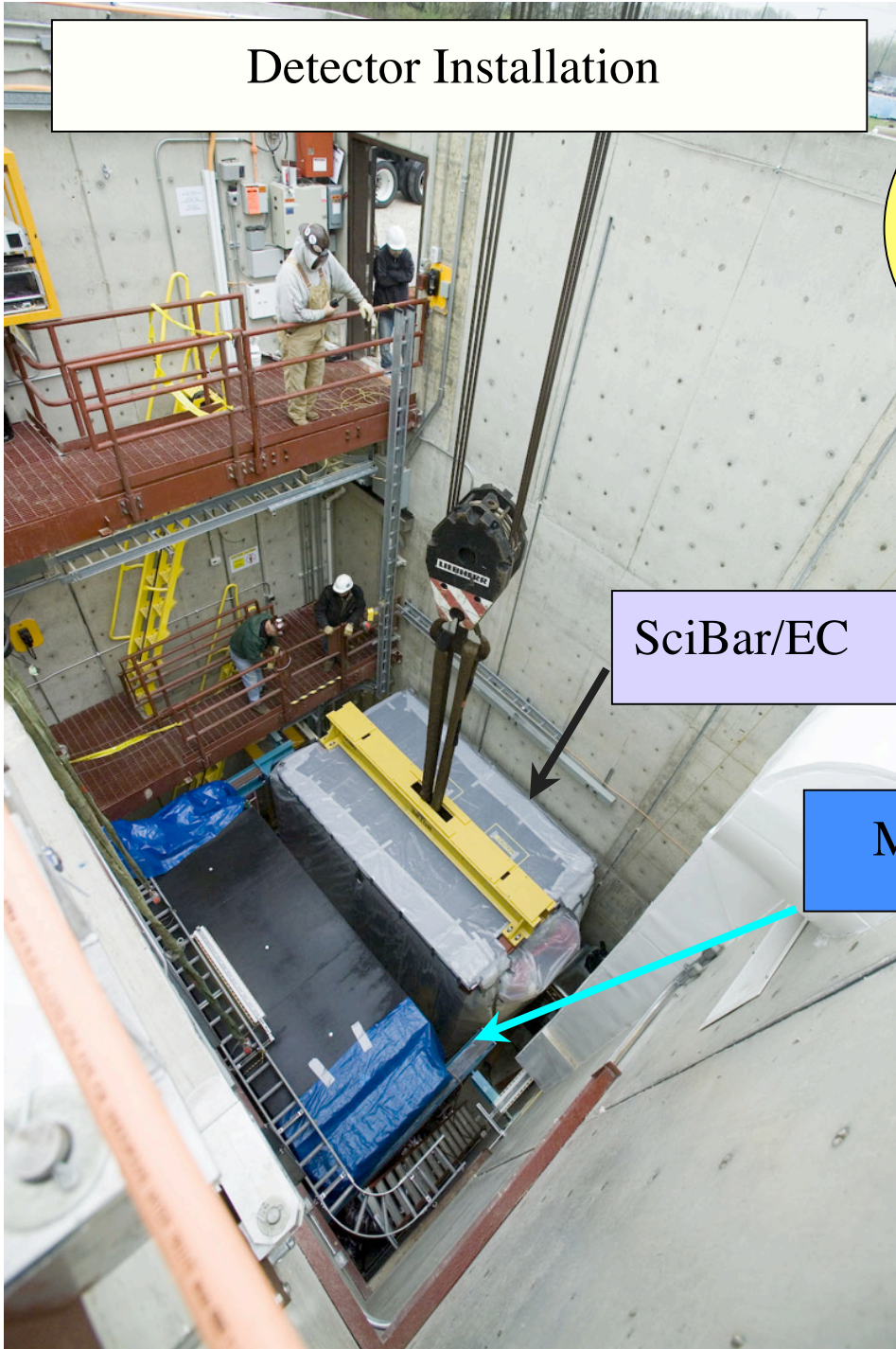
Detector Installation

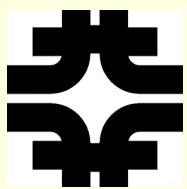
SciBooNE students worked hard to ensure the success of the installation!

SciBar/EC

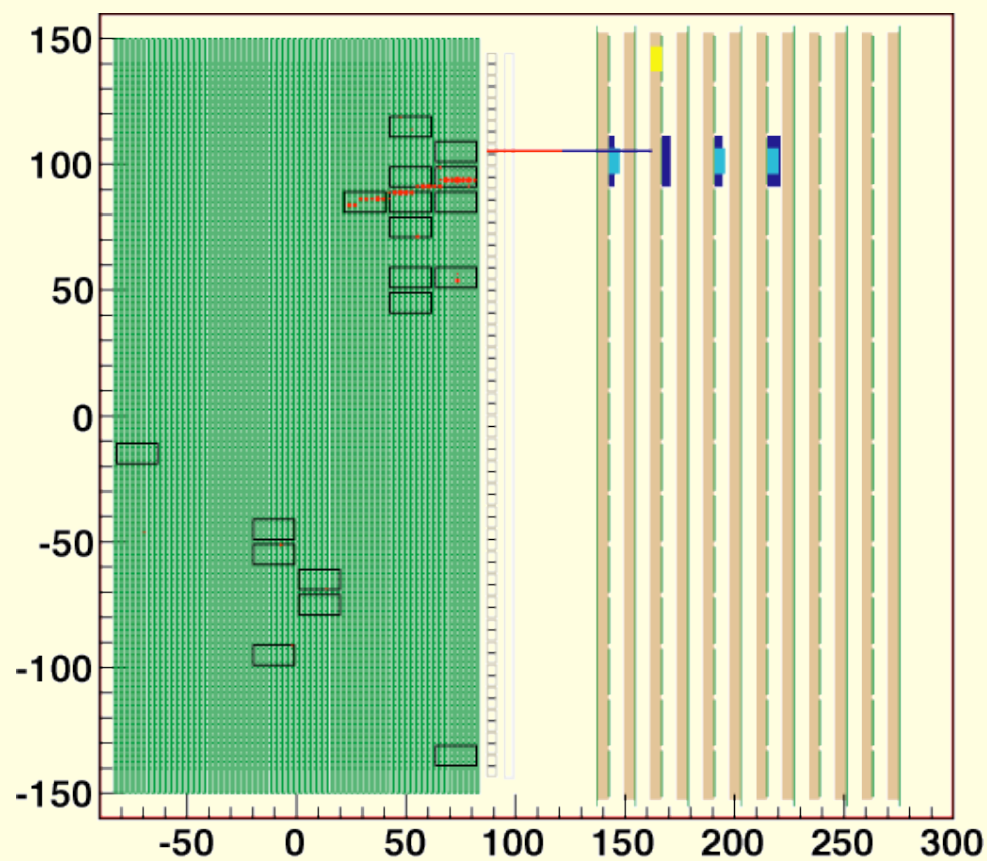
MRD

Photo of
Yasuhiro Nakajima
by Joe Walding

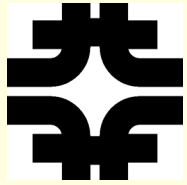




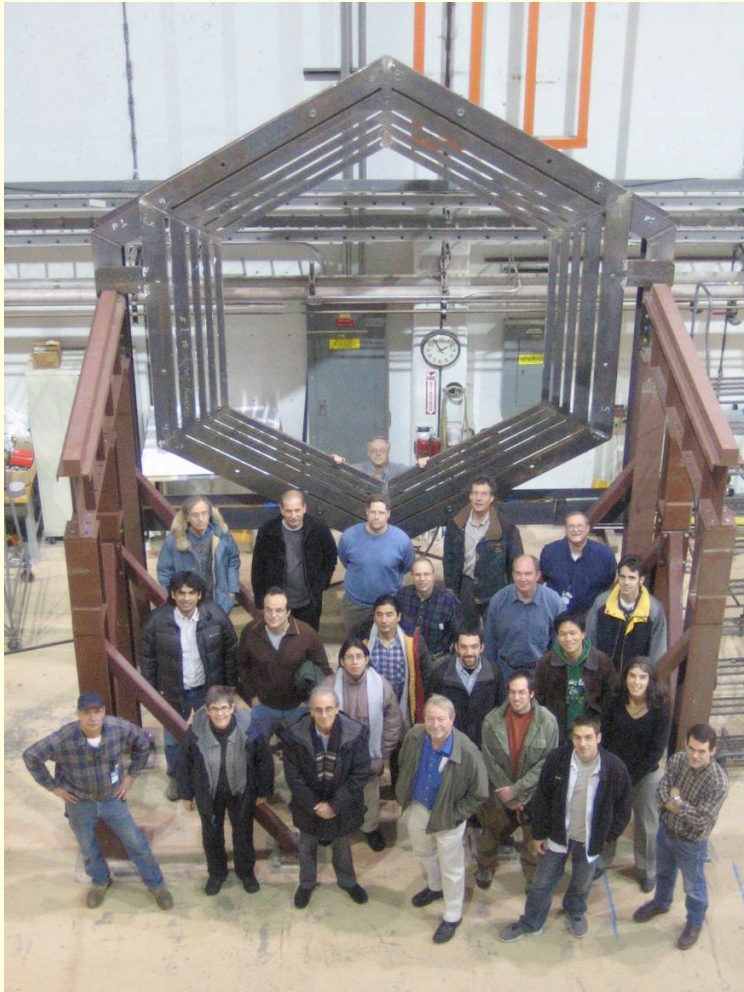
First beam neutrino event



5/28/07



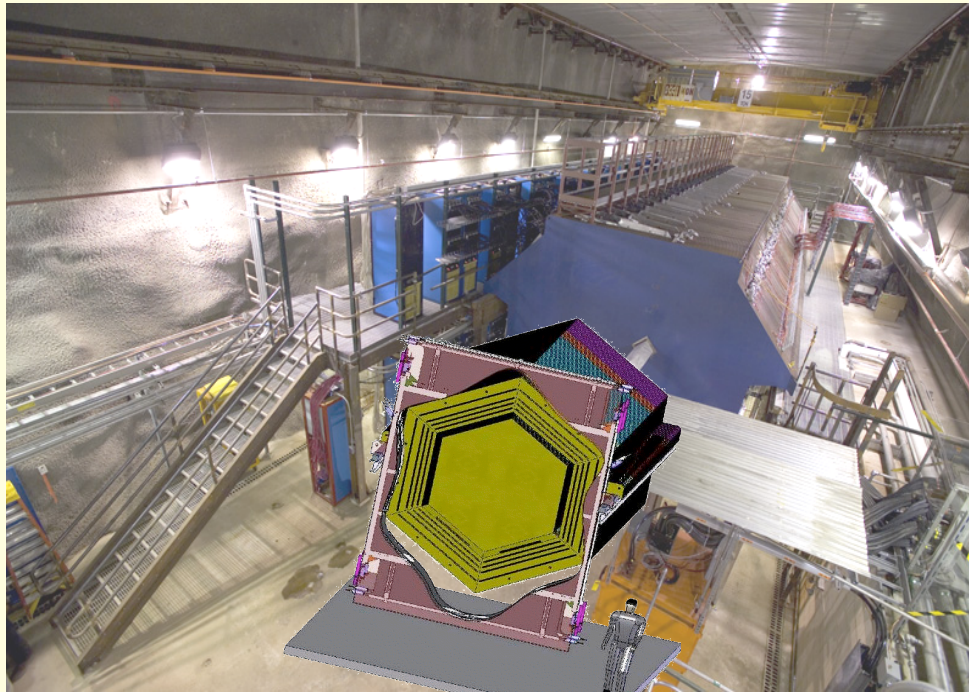
Milestones for MINERvA



- Proposed : December 2003
- PAC approval : April 2004
- CD-1/2/3a : 30 March 2007
- August 2007 : CD-3b review scheduled
- Late 2007 – early 2008: Construction of “Tracking Prototype”
 - 20 detector modules
 - Will use to study tolerance stackup, tracking, etc.
- Late 2008: Test beam detector taking data
- 2008-2009: Construction of full detector
- 2010: Installation/commissioning of detector.



MINERvA Run Plan



- 1 “year” = 4.0×10^{20} POT taking data in the LE beam configuration with MINOS
- 3 “years” = 12.0×10^{20} POT taking data in the ME beam configuration with NOvA



Event Sample:
Assume: 4.0×10^{20} POT LE and
 12.0×10^{20} POT ME beam



14.5 Million total CC events

Fiducial Volume = 3 tons CH, 0.2t He, 0.15t C, 0.7t Fe and 0.85t Pb

Expected CC event samples:

9.0 M ν events in 3 tons of CH (totally active central detector)

0.6 M ν events in He

0.4 M ν events in C

2.0 M ν events in Fe

2.5 M ν events in Pb

Main CC Physics Topics (Statistics in CH only)

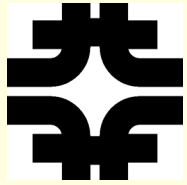
- | | |
|---|--|
| • Quasi-elastic | 0.8 M events |
| • Resonance Production | 1.7 M total |
| • Transition: Resonance to DIS | 2.1 M events |
| • DIS, Structure Funcs. and high-x PDFs | 4.3 M DIS events |
| • Coherent Pion Production | 89 K CC / 44 K NC |
| • Strange and Charm Particle Production | > 240 K fully reconstructed events |
| • Generalized Parton Distributions | order 10 K events |
| • Nuclear Effects | He: 0.6 M, C: 0.4 M, Fe: 2.0 M and Pb: 2.5 M |



News from NO ν A



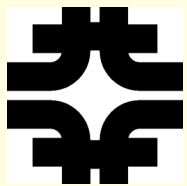
- **4/05:** PAC recommends Stage I approval for NO ν A
- **11/05:** Ray Orback approves DOE CD-0 (Mission Need) for “EvA”
- **2/06:** EvA in President’s FY07 budget for \$10.3M
- **2/06:** NuSAG report - “The U.S. Should conduct the NO ν A experiment at Fermilab”
- **2/06:** Director’s Review for CD-1 (Cost, schedule and scope range)
- **4/06:** DOE Lehman Review for CD-1 - “CD-1 Approval is recommended”
- **10/06:** P5 Report - “Proceed with the 20 kt scale NO ν A Experiment” (\$200M implied)
 - Provides information for the next step on the road to measuring CP violation, e.g. conventional beams vs neutrino factory...
- **2/07:** NO ν A in President’s FY08 budget for \$36M
- **5/02/07 :** Office of Science issues CD-1 approval
- **6/07 :** Director’s Review in preparation for CD-2/3a review
- **7/07 :** DOE CD-2/3a review scheduled



The NOvA Project



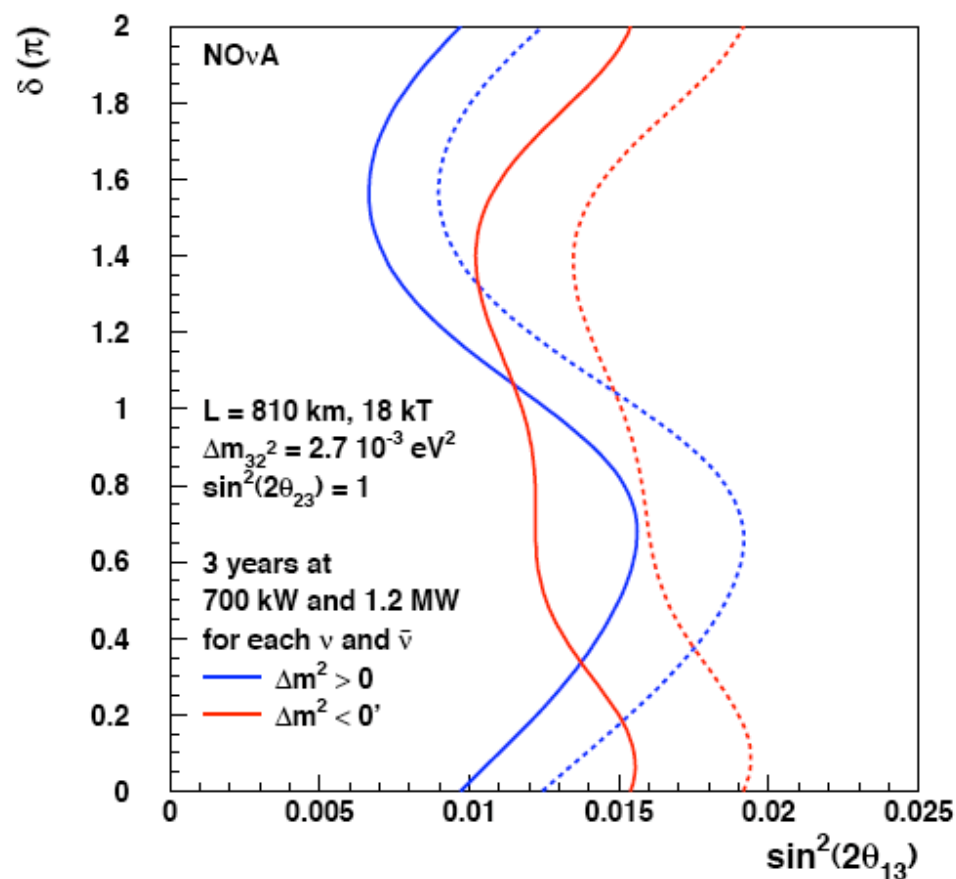
- Liquid scintillator in 3.9 cm (W) x 6.6 cm (D) x 15.7 m (L) cells
 - ~20 kton (goal) but TPC is “capped” @ \$260 M
 - ~500,000 cells in alternating x-y planes
 - 4 million gallons of liquid scintillator
 - 17,000 km of fiber



Physics with NOvA



3σ Sensitivity to $\theta_{13} \neq 0$

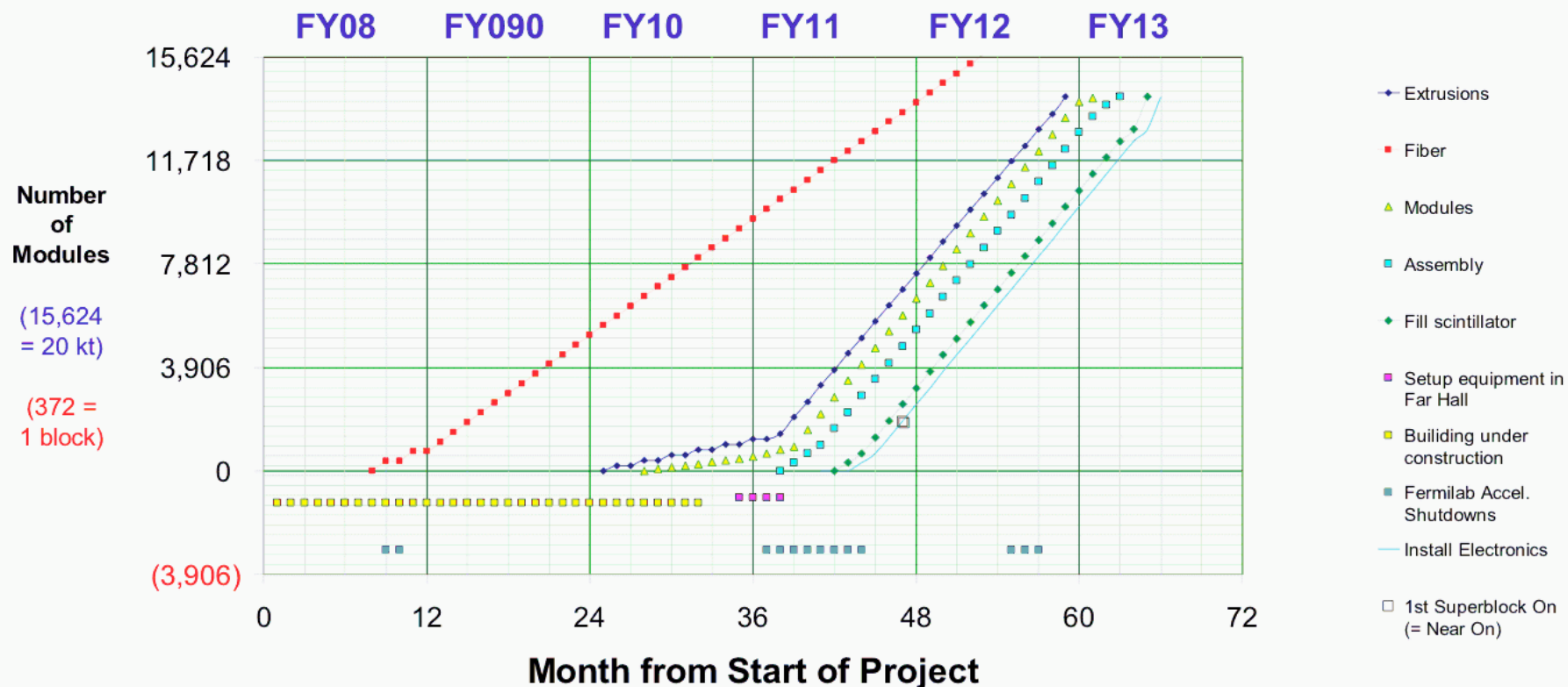


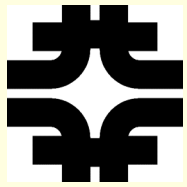
- Primary Goal :
 - Study of $\nu_\mu \rightarrow \nu_e$ oscillations
 - $\sin^2 2\theta_{13}$
 - the sign of Δm_{32}^2
 - the CP-violating phase δ
- Level of “success” depends on *NATURE*
 - Current limit on $\sin^2 2\theta_{13} \sim 0.14$ (90% CL)



Schedule

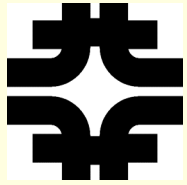
- Critical path is building, then assembly at Ash River





θ_{13} Program

- Phase I (~2009 - ~2013):
 - World wide effort to measure θ_{13}
 - Reactor experiments :
 - Double Chooz and Daya Bay
 - Accelerator long baseline :
 - JPARC to SuperK (T2K) and NO ν A
 - Phase I experiments will determine if $\sin^2 2\theta_{13} > 0.05$ very quickly
 - If **$\sin^2 2\theta_{13} > 0.05$** NO ν A may be able to determine the mass hierarchy (depends on the value of δ_{CP})
 - If **$0.02 < \sin^2 2\theta_{13} < 0.05$** it will take longer to determine this;
 - determining the mass hierarchy will be difficult and sensitivity to CP is achievable but even more challenging



Future Long Baseline Study (sponsored by BNL/FNAL)

March 2006 -
May 2007

arXiv:0705.4396
(May 2007)

Fermilab-0801-AD-E
BNL-77973-2007-IR

Report of the US long baseline neutrino experiment study

V. Barger,¹ M. Bishai,² D. Bogert,³ C. Bromberg,⁴ A. Curioni,⁵ M. Dierckxsens,²
M. Diwan,² F. Dufour,⁶ D. Finley,³ B. T. Fleming,⁵ J. Gallardo,² D. Gerstle,⁵ J. Heim,²
P. Huber,¹ H. Jostlein,³ C. K. Jung,⁷ S. Kahn,² E. Kearns,⁶ H. Kirk,² T. Kirk,⁸ K. Lande,⁹
C. Laughton,³ W. Y. Lee,¹⁰ K. Lesko,¹⁰ C. Lewis,¹¹ P. Litchfield,¹² A. K. Mann,⁹
A. Marchionni,³ W. Marciano,² D. Marfatia,¹³ A. D. Marino,³ M. Marshak,¹² S. Menary,¹⁴
K. McDonald,¹⁵ M. Messier,¹⁶ W. Pariseau,¹⁷ Z. Parsa,² S. Pordes,³ R. Potenza,¹⁸ R. Rameika,³
N. Saoulidou,³ N. Simos,² R. Van Berg,⁹ B. Viren,² W.T. Weng,² K. Whisnant,¹⁹ R. Wilson,²⁰
W. Winter,²¹ C. Yanagisawa,⁷ F. Yumiceva,²² E. D. Zimmerman,⁸ and R. Zwaska³

¹Department of Physics, University of Wisconsin, Madison, WI 53706, USA

²Physics Department, Brookhaven National Laboratory, Upton, NY 11973, USA

³Fermi National Accelerator Laboratory, Batavia, IL 60510, USA

⁴Department of Physics and Astronomy,

Michigan State University, East Lansing, MI 48824, USA

⁵Department of Physics, Yale University, New Haven, CT 06520, USA

⁶Department of Physics, Boston University, Boston, MA 02215, USA

⁷Stony Brook University, Department of Physics and Astronomy, Stony Brook, NY 11794, USA

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⁹Department of Physics and Astronomy,

University of Pennsylvania, Philadelphia, PA 19104, USA

¹⁰Lawrence Berkeley National Laboratory,

Physics Division, Berkeley, CA 94720, USA

¹¹Department of Physics, Columbia University, New York, NY 10027, USA

¹²School of Physics and Astronomy, University of Minnesota, Minneapolis, MN 55455, USA

¹³Department of Physics and Astronomy,

University of Kansas, Lawrence, KS 66045, USA

¹⁴Department of Physics and Astronomy,

York University, Toronto, Ontario M3J1P3, Canada

¹⁵Department of Physics, Princeton University, Princeton, NJ 08544, USA

¹⁶Department of Physics, Indiana University, Bloomington, IN 47405, USA

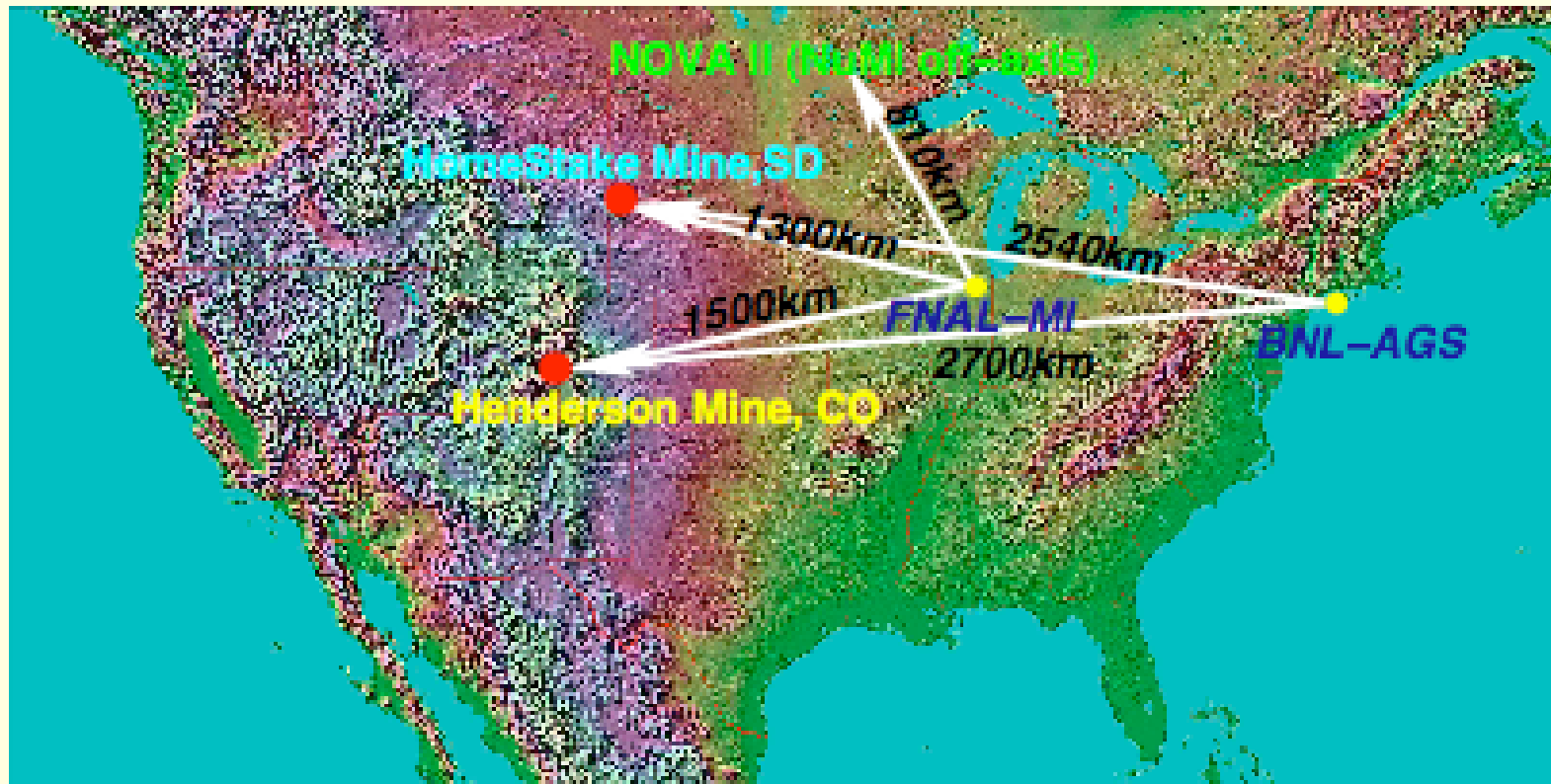
¹⁷Department of Mining Engineering,

University of Utah, Salt Lake City, UT 84112, USA

Special thanks to Milind Diwan, BNL

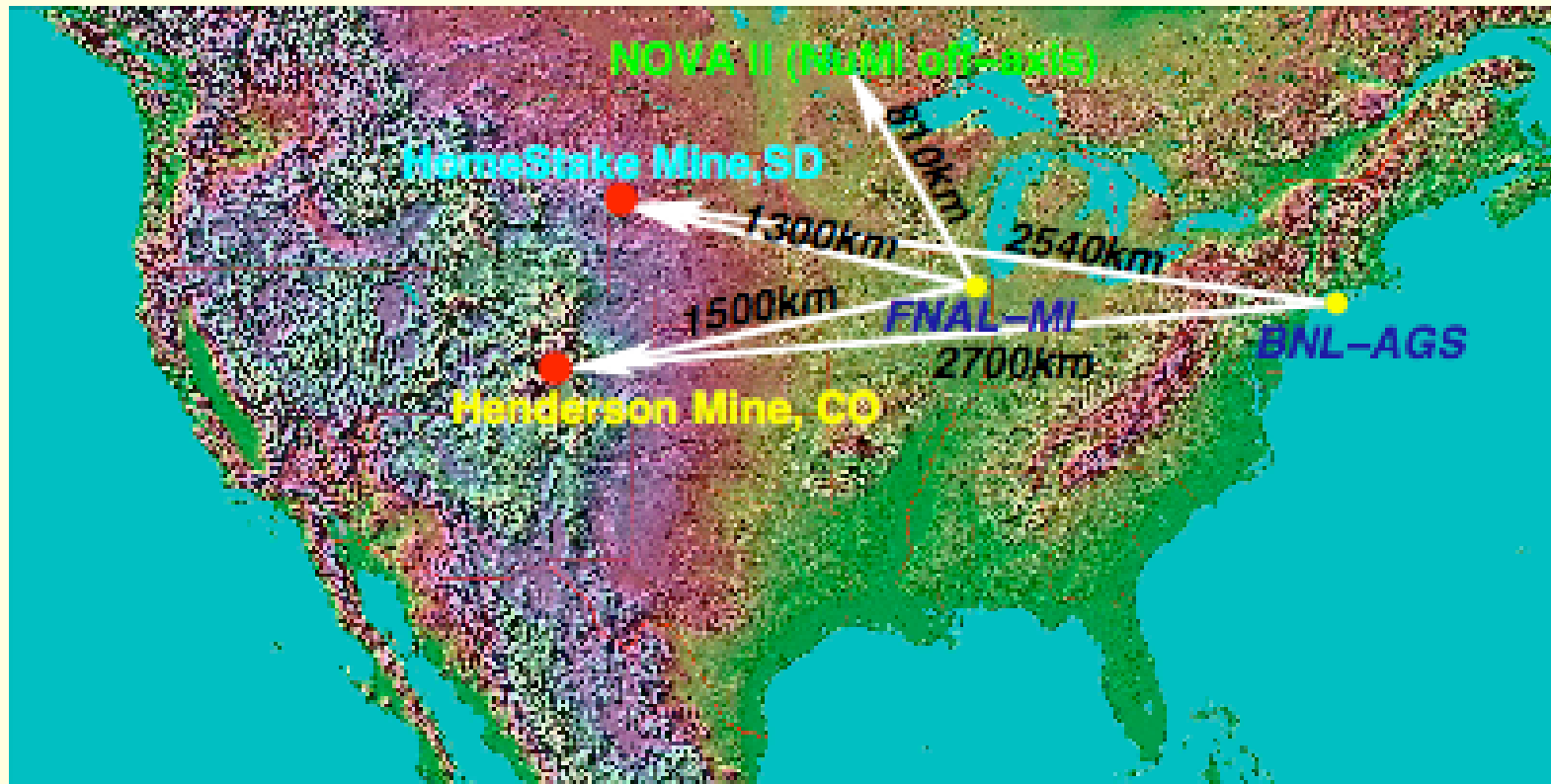


Future Long Baseline Options

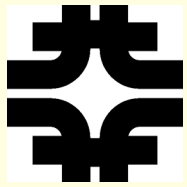




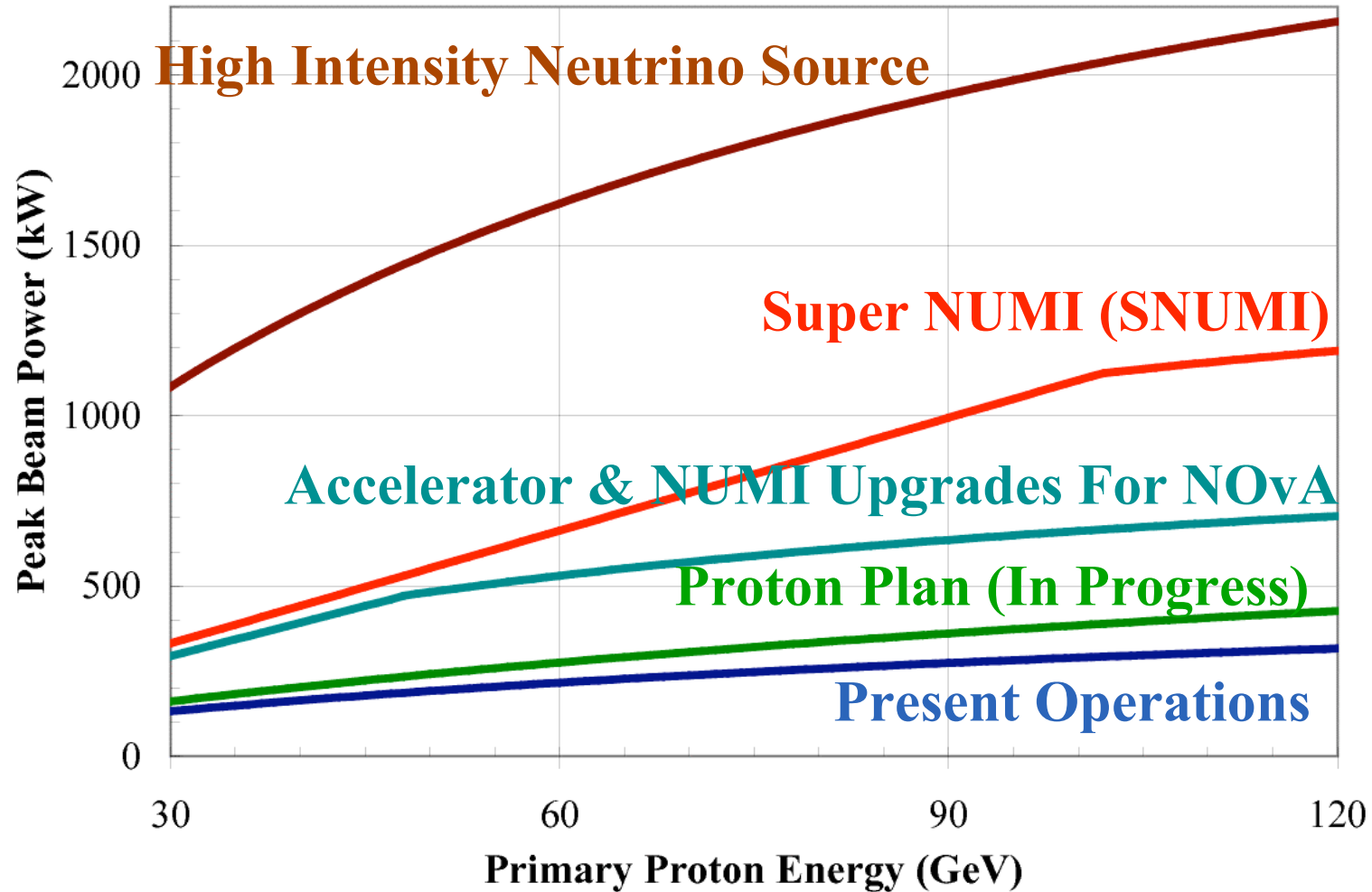
Future Long Baseline Options

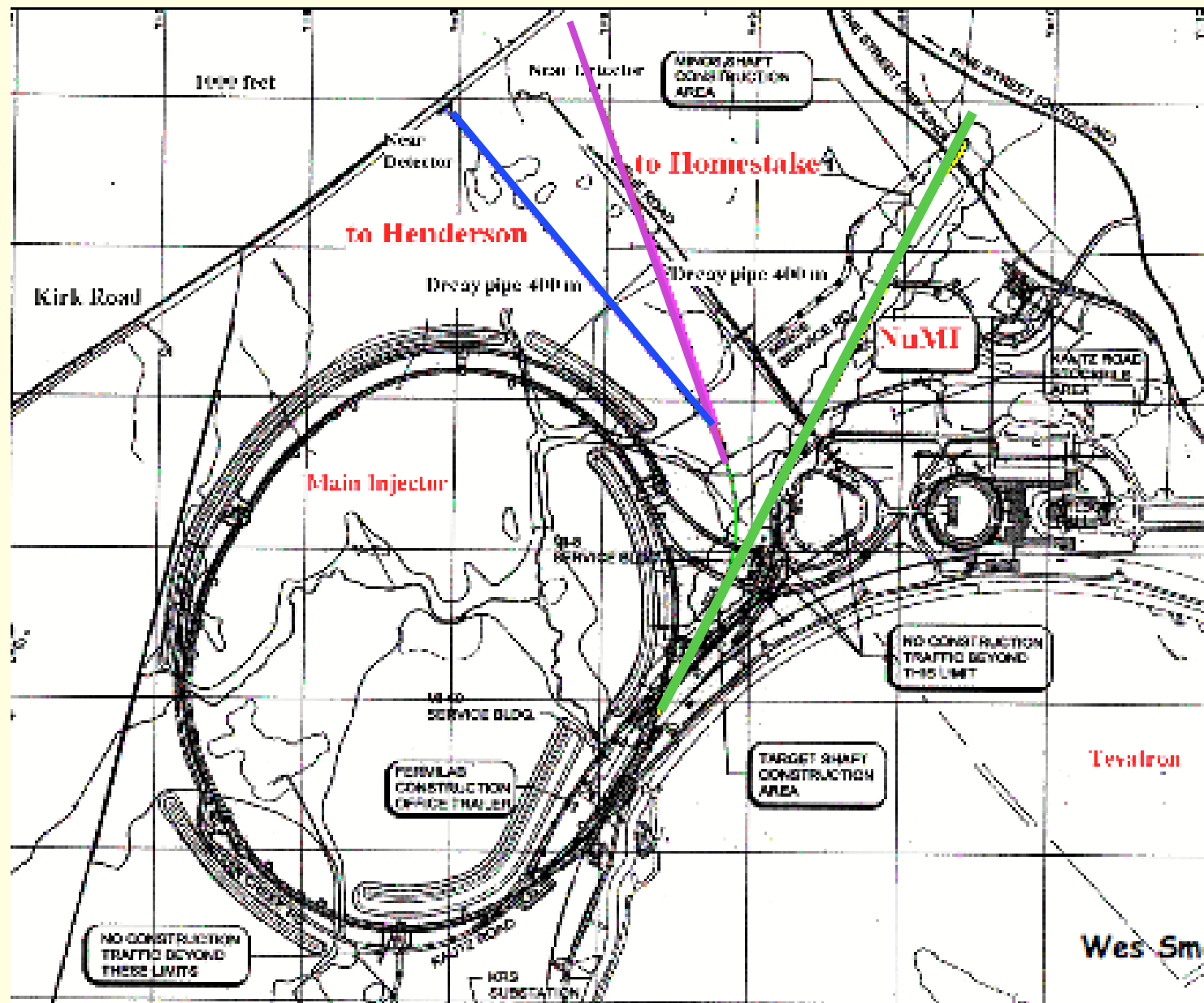


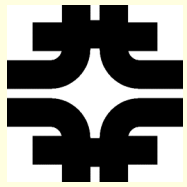
Study focused only on the option of using the Fermilab Accelerator complex.



Options for Proton Power

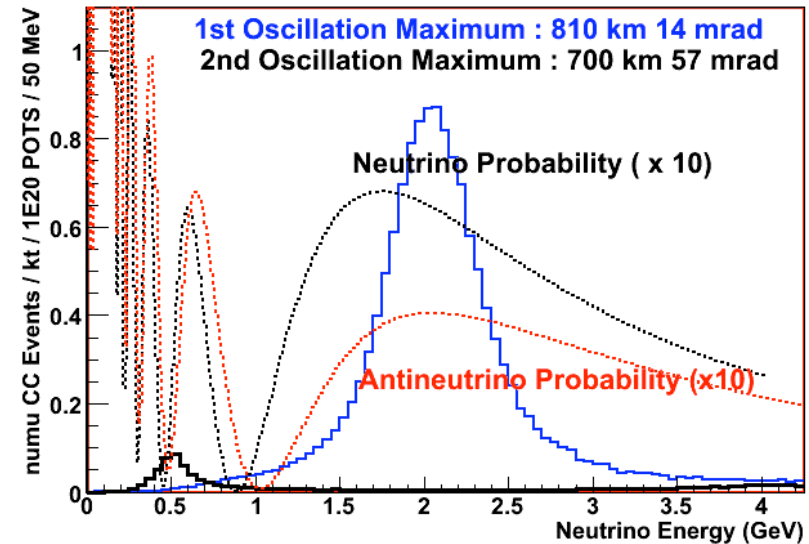
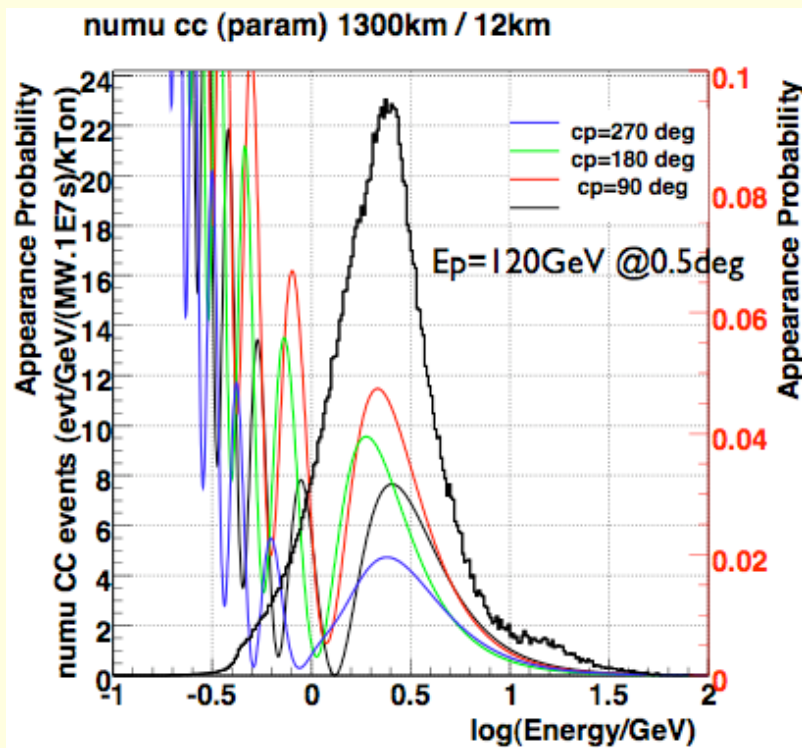






ν beam Options

- NuMI - Off-Axis (narrow band)



- NEW - Wide Band Beam

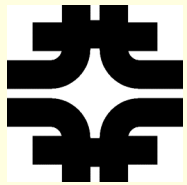


Some conclusions of the study

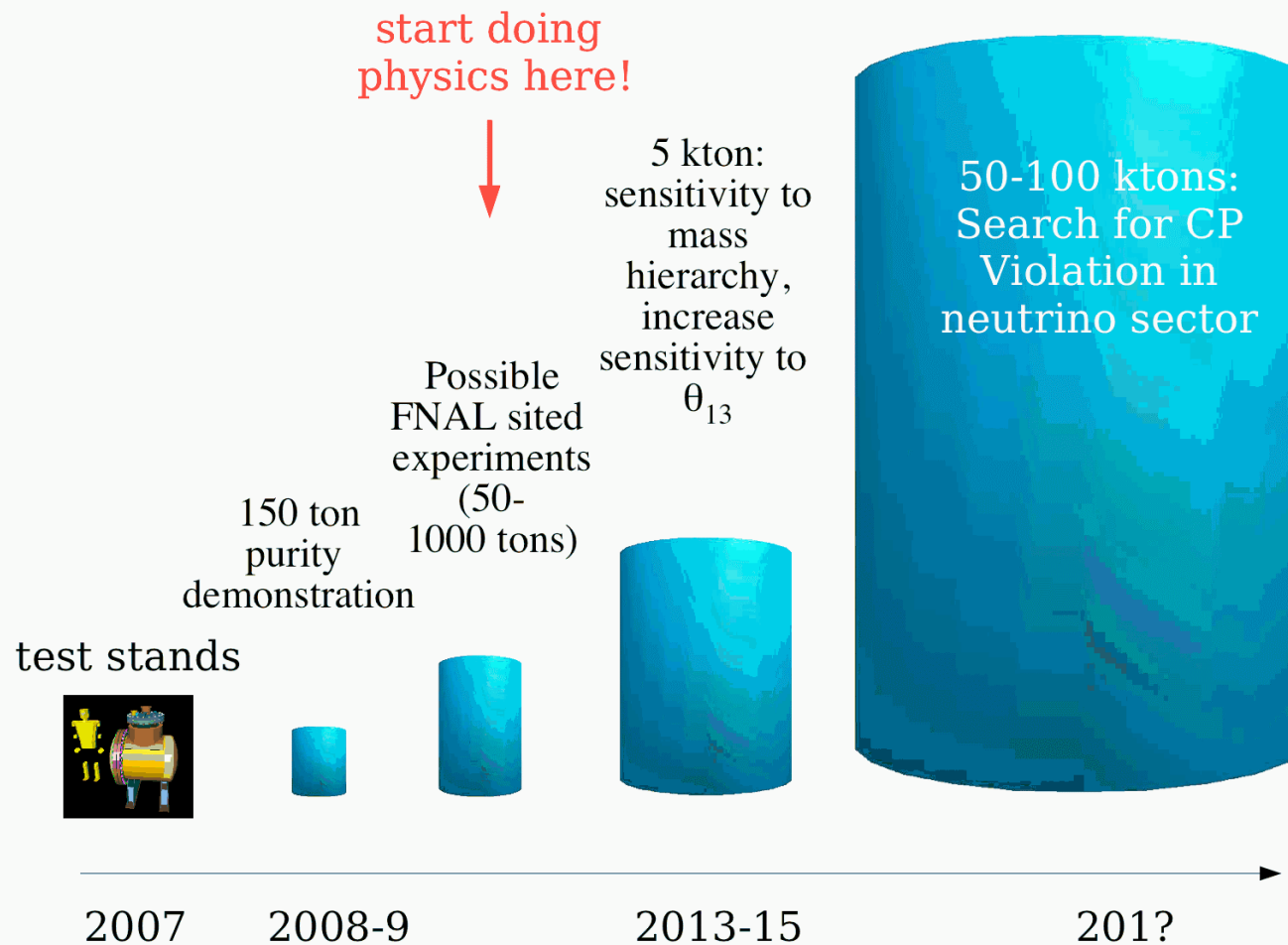
- Two Options for a future program - how to choose will be easier once Phase I experiments present results
 - **Existing NuMI Off-axis** option is good for
 - $\sin^2 2\theta_{13} \neq 0$ as low as 0.005
 - Mass hierarchy for large “ish” θ_{13}
 - Specific sensitivity depends on protons, mass and running time
 - possible exploration of CP violation in neutrino sector
 - **New WBB at a longer baseline** may be necessary for determining the mass hierarchy for all values of δ_{CP} if $\sin^2 2\theta_{13}$ turns out to be quite small
 - Win with baseline and the Wide band beam which enables one to fit the observed spectrum to the oscillation parameters



- Either option requires
 - Detector masses $\gg 100\text{kton}$
 - Megawatt class proton source
- R &D on large scale detectors necessary
 - Water cerenkov
 - Proven, well understood technology
 - Scaling to $\sim 100\text{-}300\text{ kton}$ seems feasible
 - Needs to operate underground
 - Liquid argon
 - Potentially high resolution and excellent background rejection
 - Cost and Technical scaling is a challenge
 - Possibly could operate on the surface

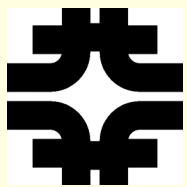


Concept for evolution

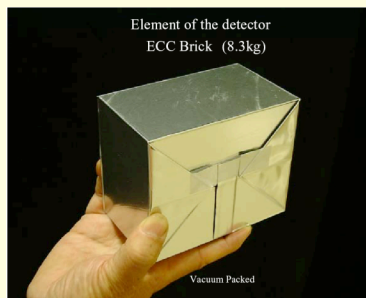


http://www.fnal.gov/directorate/Longrange/Steering_Public/documents.html

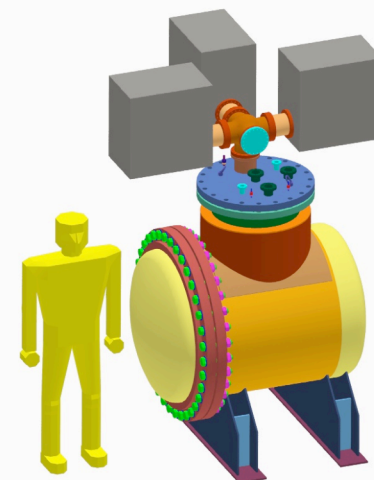
B. Fleming
5/31/07



Detector R & D

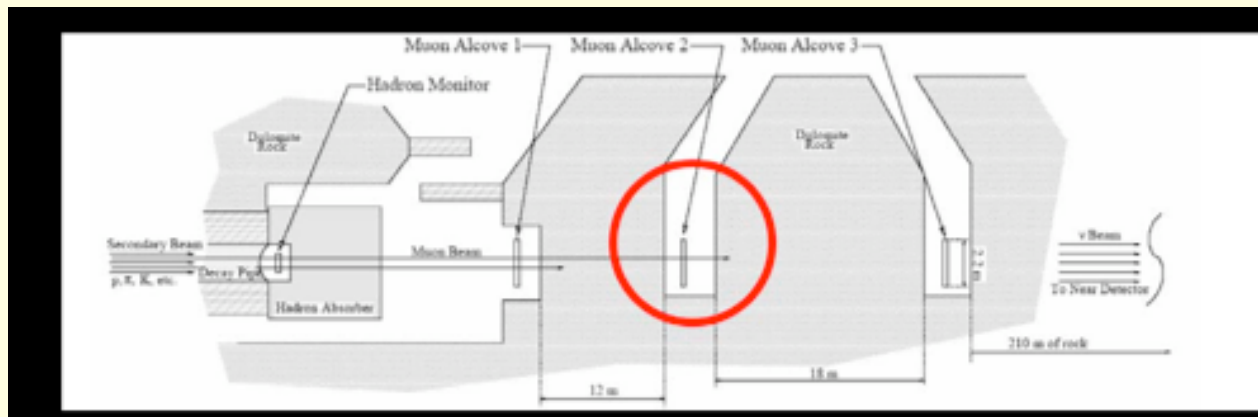


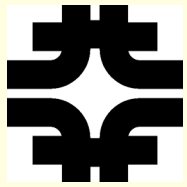
T-952
OPERA emulsion
bricks



T-962 mini-
LAr TPC

T-968 T2K MUMON





Conclusions & Outlook

- Fermilab neutrino program is active and diverse.
- Operating experiments **MINOS** and **MiniBooNE** are producing exciting results and publications.
- **SciBooNE** has completed installation and is commissioning with neutrino beam
- The **MINERvA** experiment is on the approval path, and progressing well in R&D and prototyping
- The **NOvA** project is also progressing through the review process and will hopefully move from R&D to construction before the end of this FY.
- Planning for a θ_{13} **Phase II experiment** is an important part of our neutrino program
- Our existing facilities are being used to support a variety of **R&D test activities** for both the Fermilab and the world wide neutrino program.